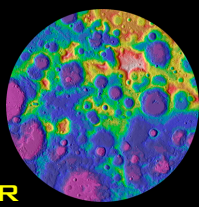


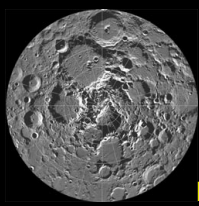
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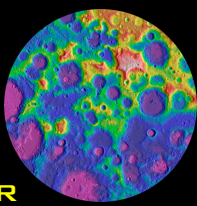
Bullialdus Crater: Probing Mineralogy and Local Hydroxyl Abundance



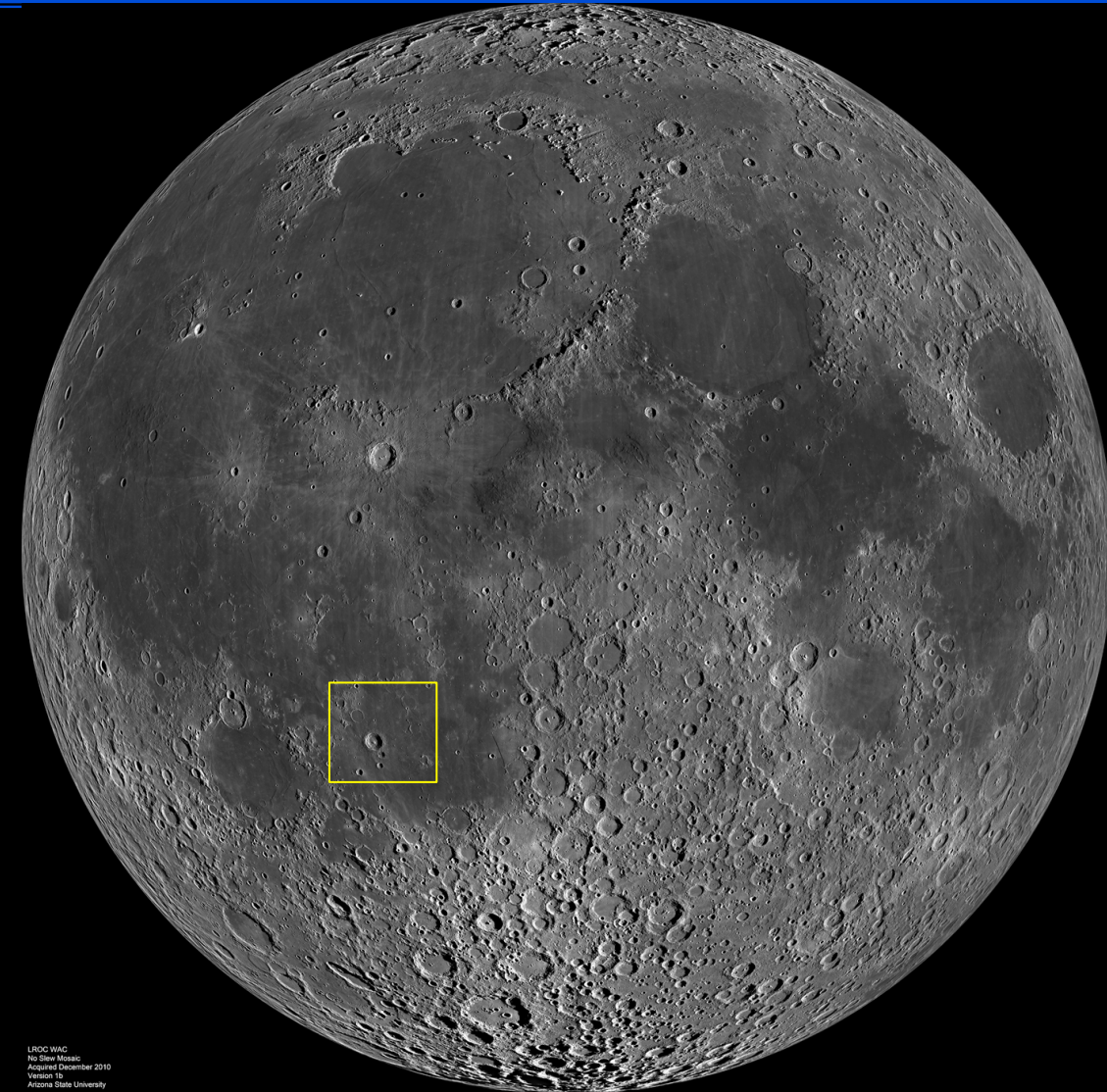
Rachel Klima, Joshua Cahill,
Justin Hagerty and David Lawrence
Lunar Science Forum, 2012
July 18, 2012



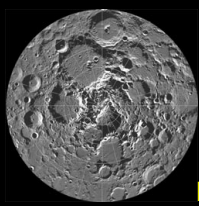
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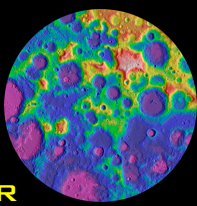
Bullialdus Crater



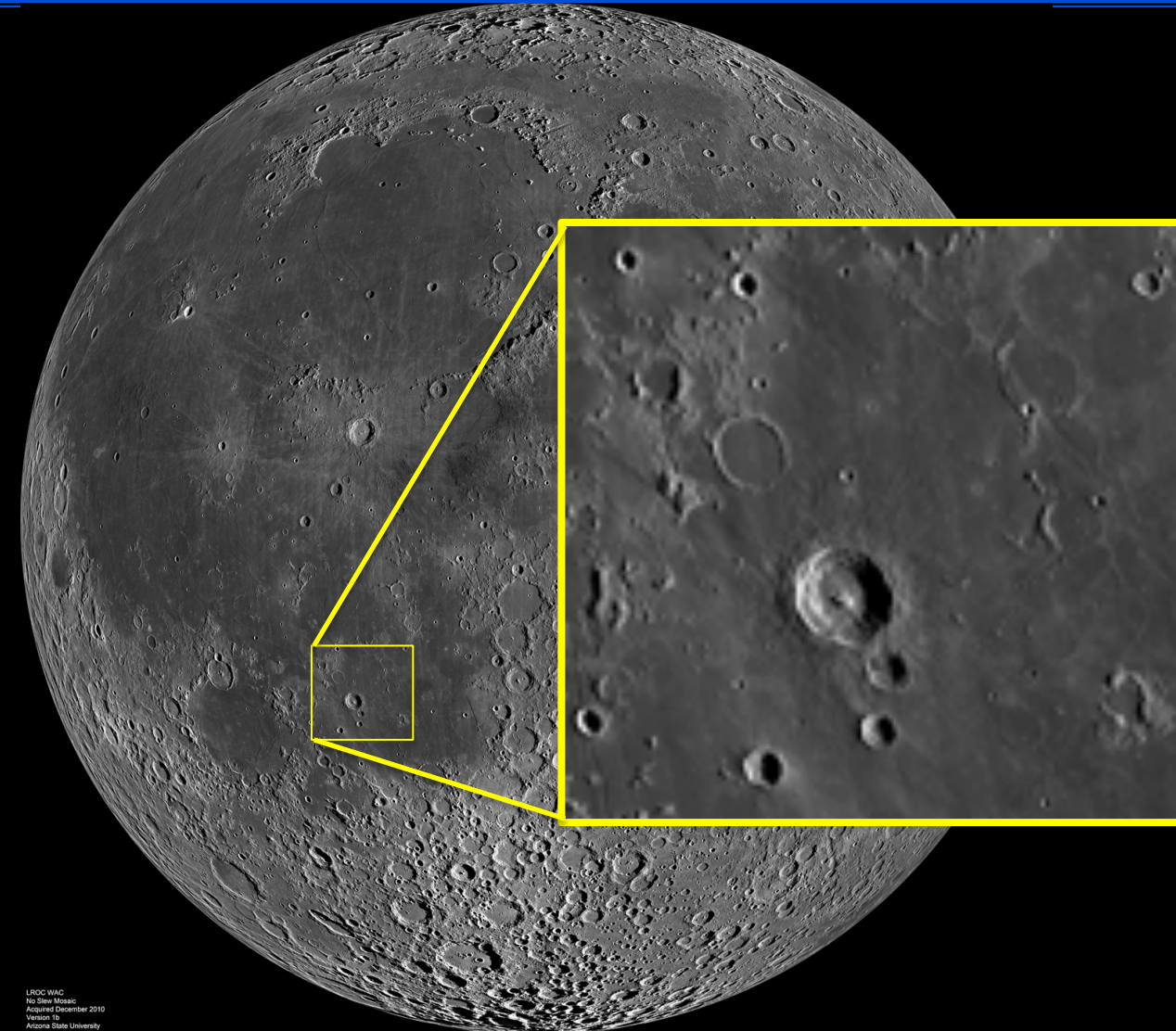
- 20.7°S, 22.2°W on western edge of Mare Nubium
- 61 km diameter
- Eratosthenian-aged



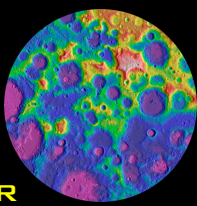
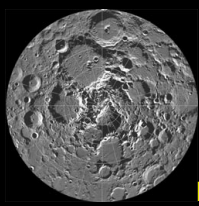
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Bullialdus Crater



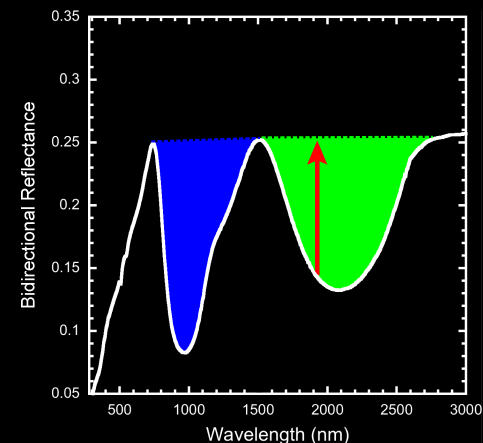
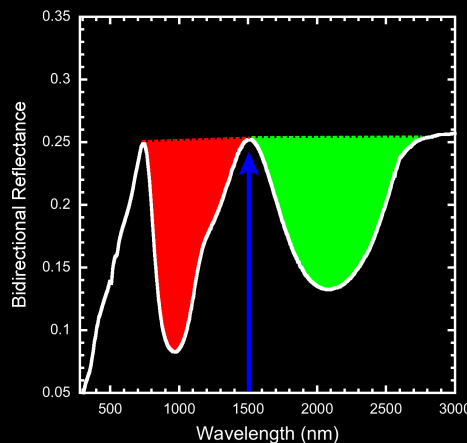
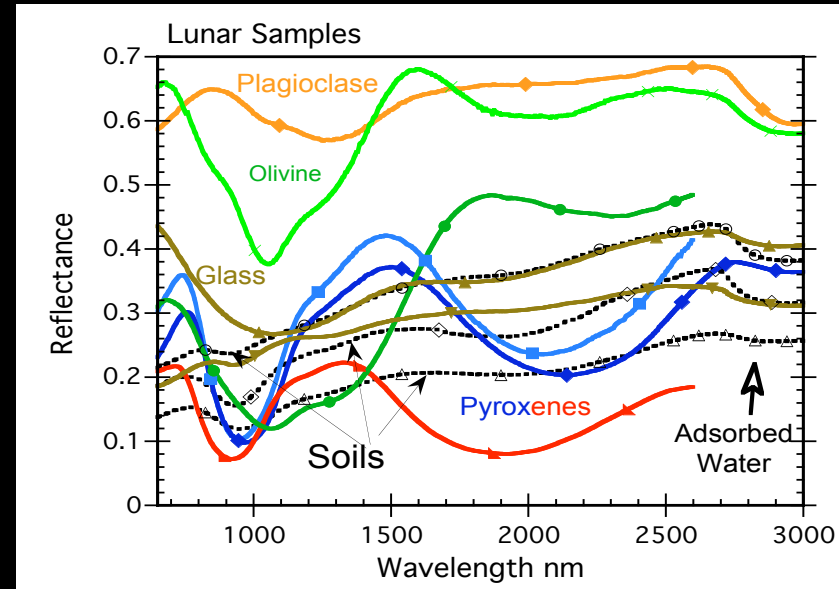
- Central peak nortic, walls more enhanced in clinopyroxene
- Layered mafic pluton (Pieters, 1991)
- Other possibilities include impact through thin old basalt flows, or through a differentiated melt sheet from the Nubium impact (Tompkins et al., 1994).
- Radiative transfer modeling suggests immature regions of the peak range from anorthositic norite, anorthositic gabbro-norite to norite (Cahill and Lucey, 2007)
- Mean $Mg' 70$ (Cahill et al. 2009)
- Norites in central peak modeled to be $>Mg'_{75}$ (Klima et al. 2011).

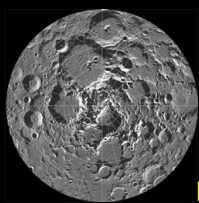


M³ Data: Color Composites



- M³ standard color composite:
Red = Integrated 1 μm band depth
Green = Integrated 2 μm band depth
Blue = Reflectance at 1.58 μm
Highlands Blue, Pyroxene Yellow
- Pyroxene composite:
Red = 1.9 μm band depth
Green = Integrated 2 μm band depth
Blue = Integrated 1 μm band depth
Highlands Black, LCP Yellow, HCP Cyan

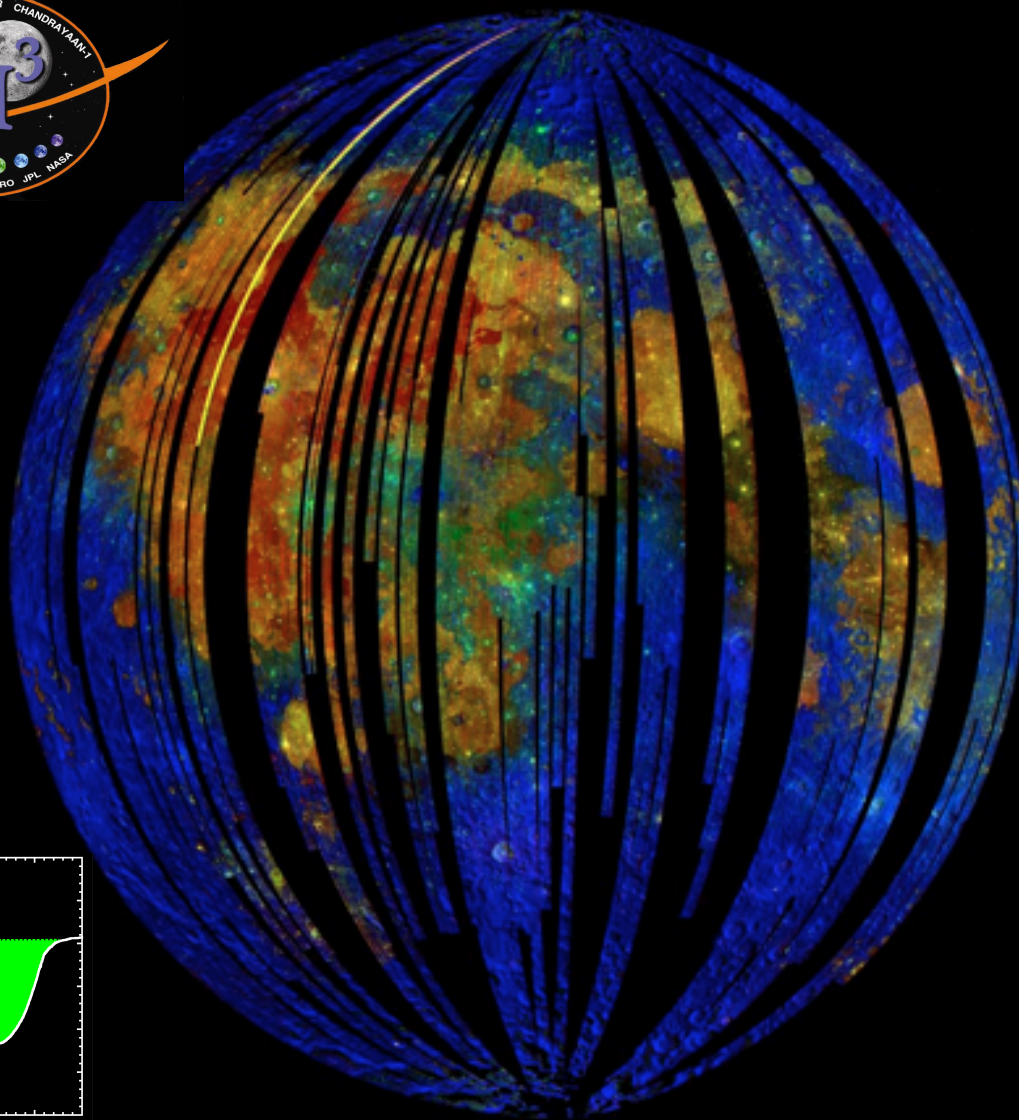




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Lunar Mineralogy: A Hyperspectral View



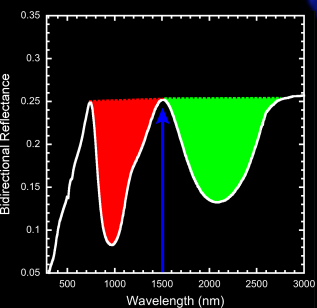
Moon Mineralogy Mapper
(M³)
Vis-NIR Imaging
Spectrometer

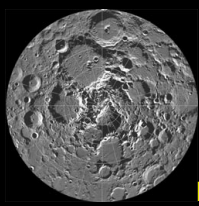
Global Mode:
85 bands

20-40 nm spectral
sampling

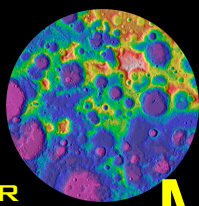
140 m/pixel at 100 km
orbit

0.4-3 μ m wavelength
range

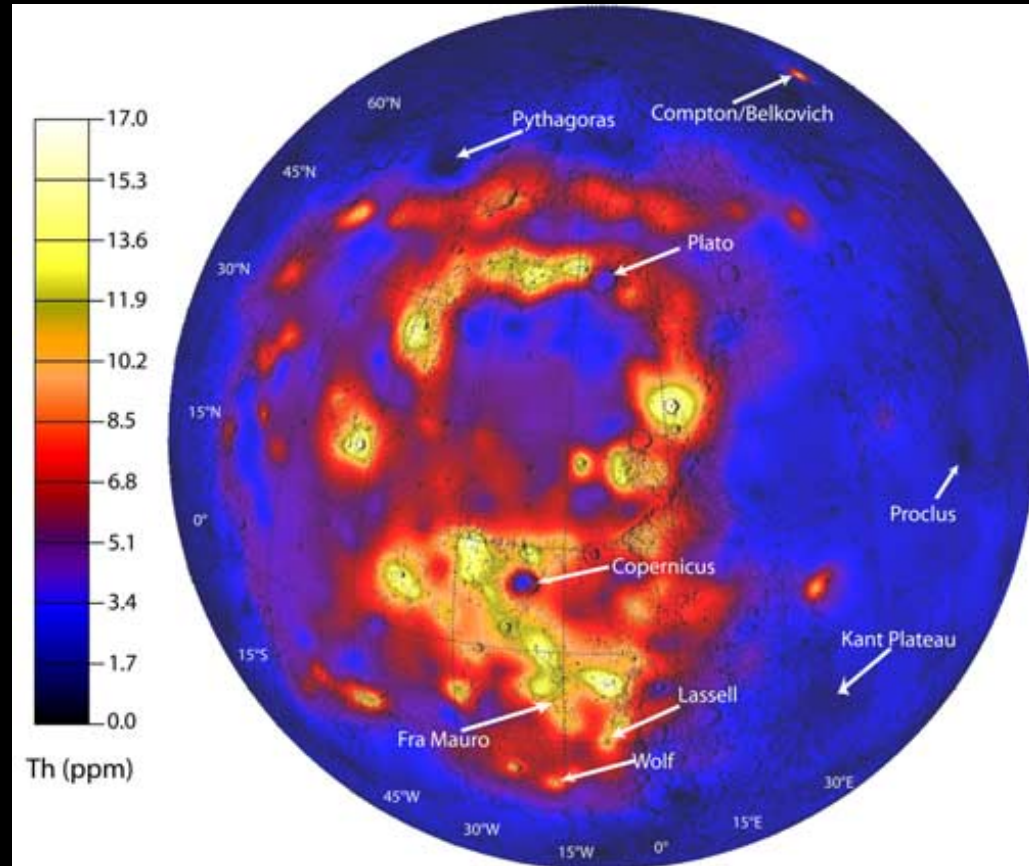
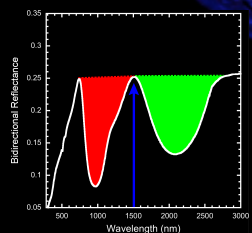
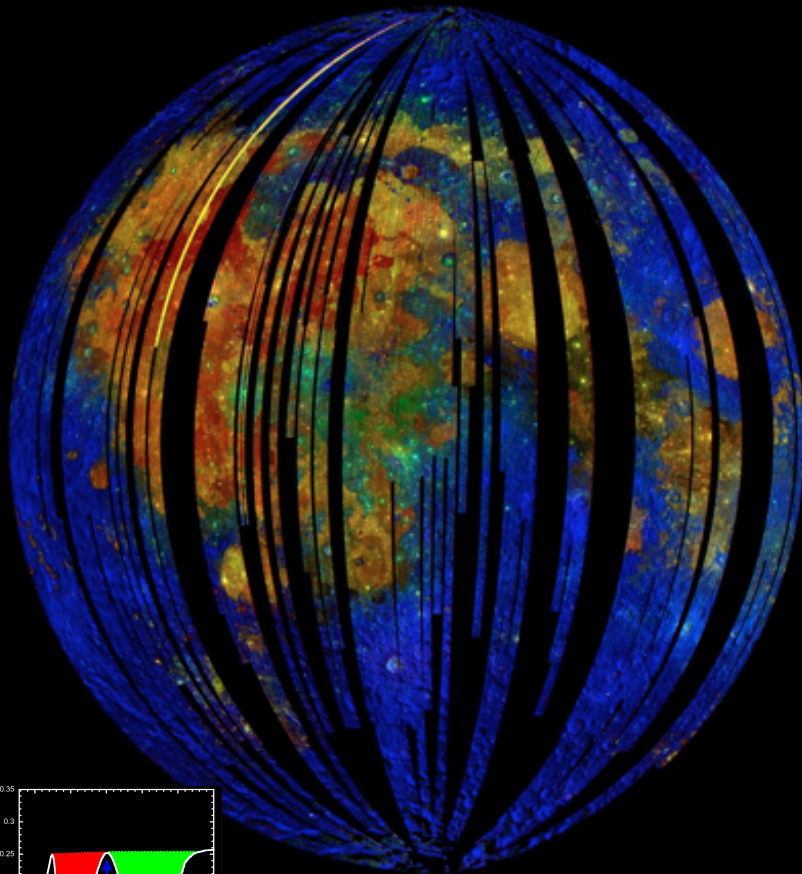


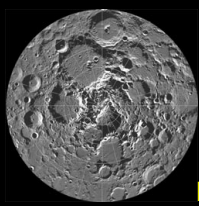


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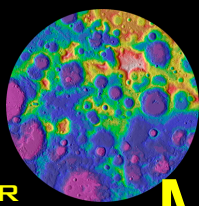


Southern PKT – Mineralogy and Thorium

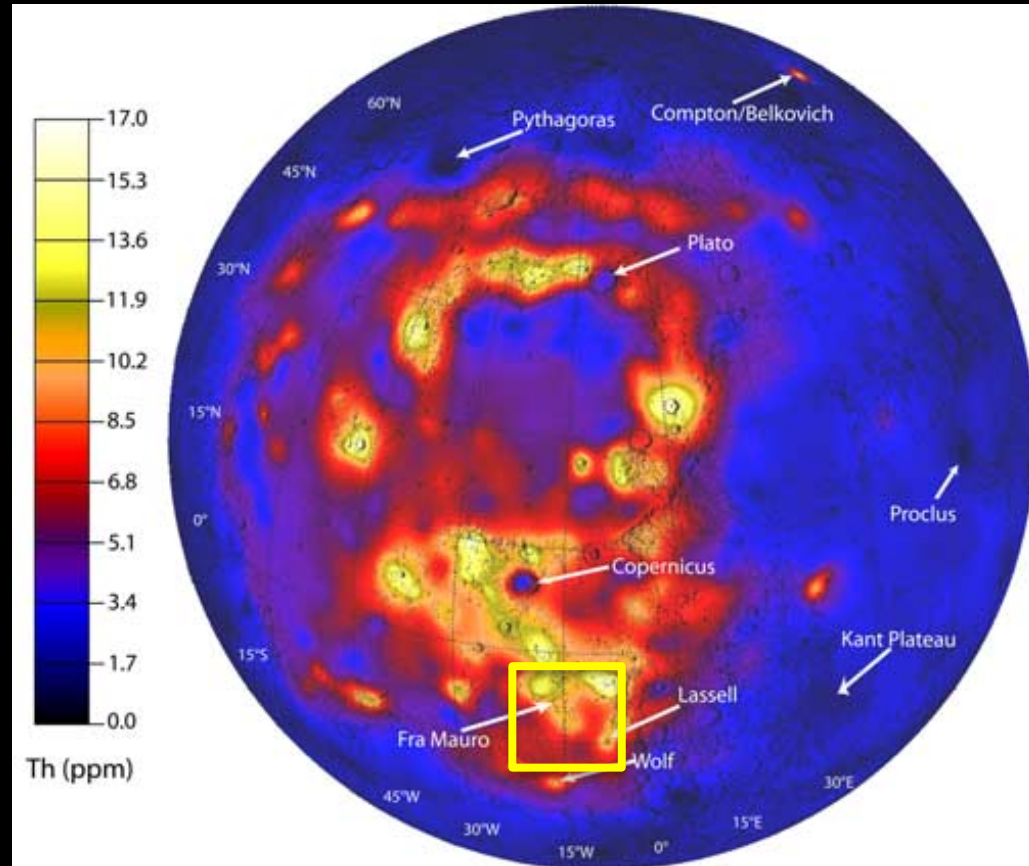
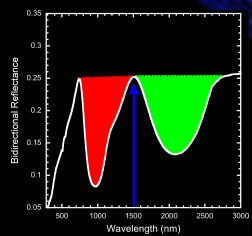
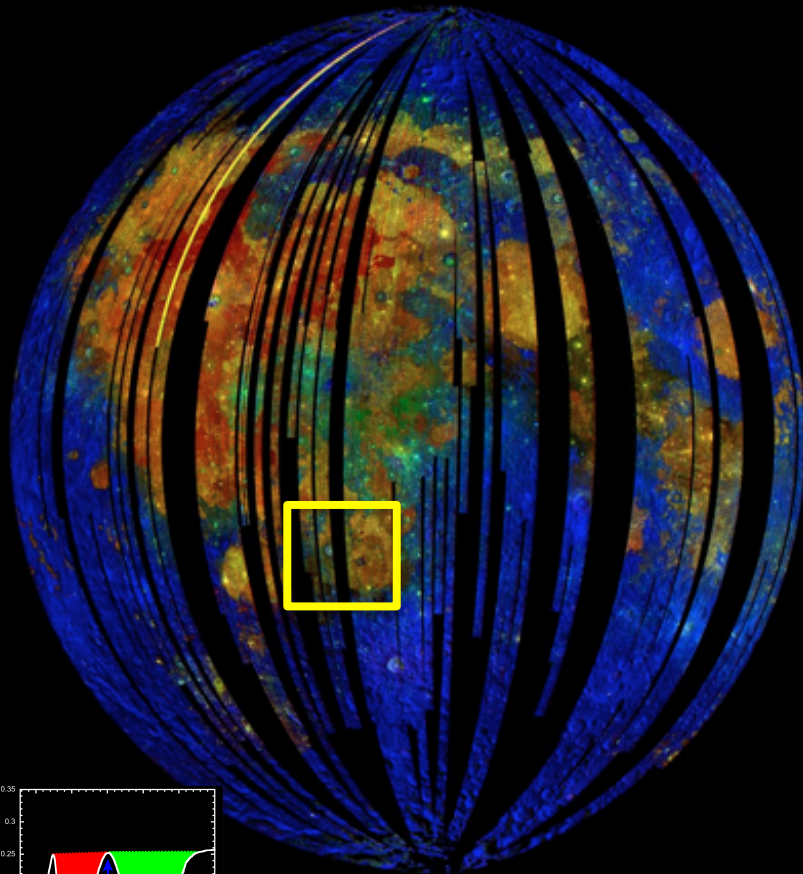


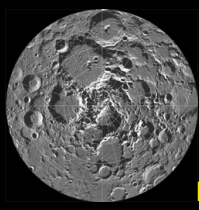


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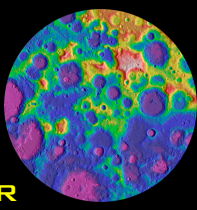


Southern PKT – Mineralogy and Thorium





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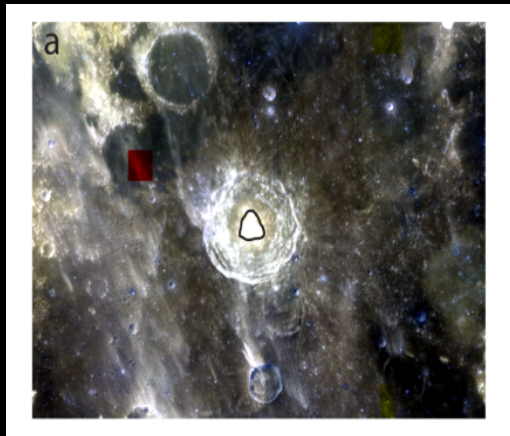


Bullialdus Region: Thorium

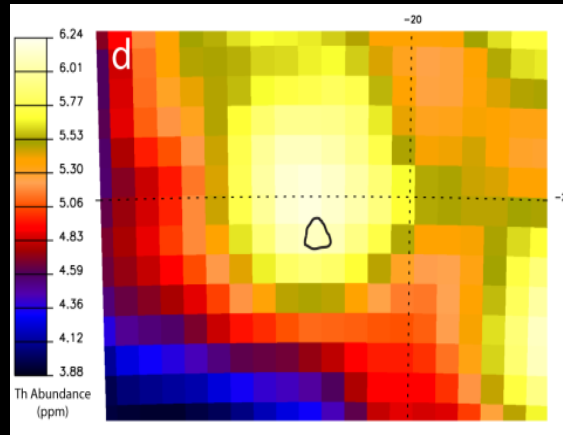


Enhanced Thorium detected at Bullialdus
corresponds to anorthositic and noritic terrain

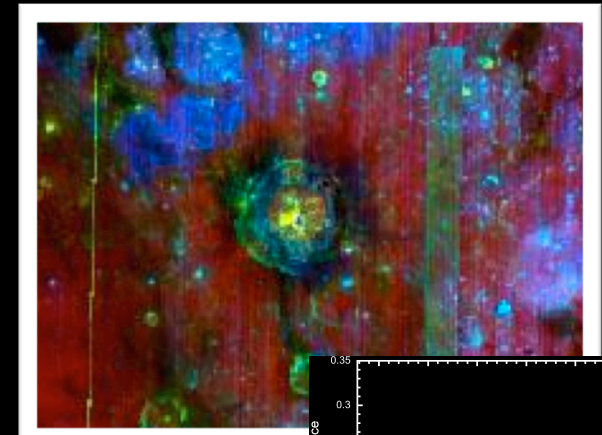
Clementine



Lunar Prospector Thorium



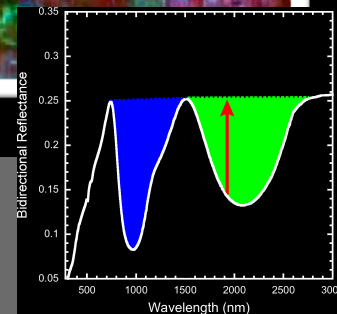
M³

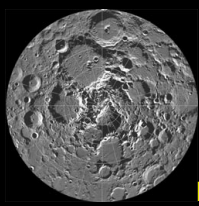


M³ Color Composite:

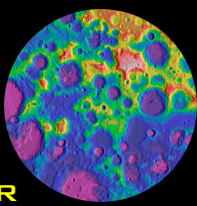
Yellow: Low-calcium pyroxene Black: Anorthosite

Blue/Cyan/Purple: Mare basalt or enriched in high-calcium pyroxene





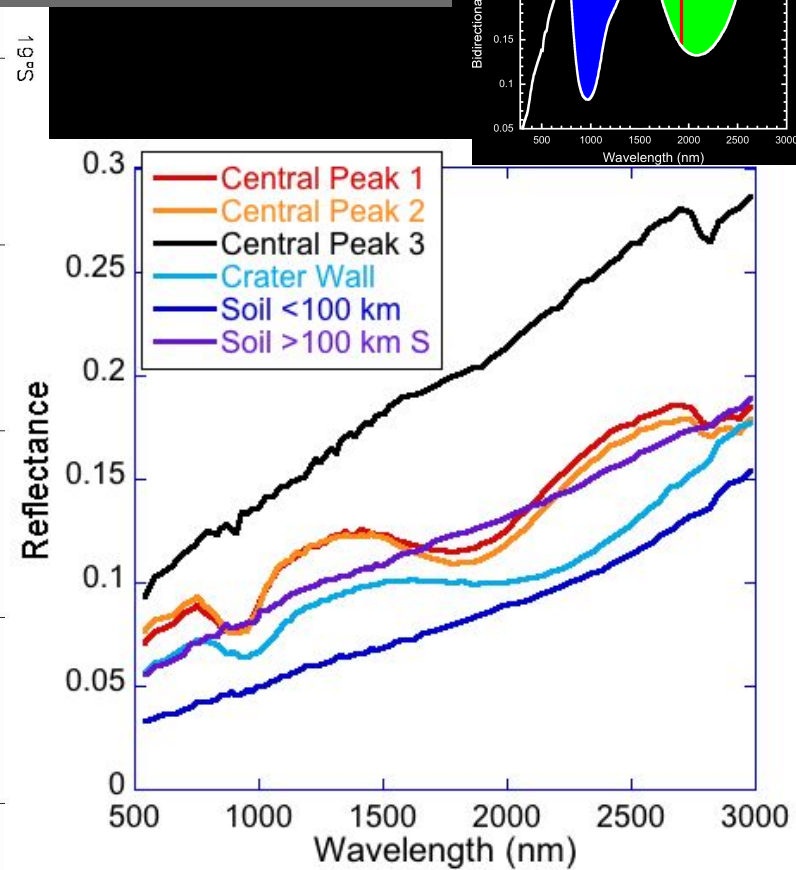
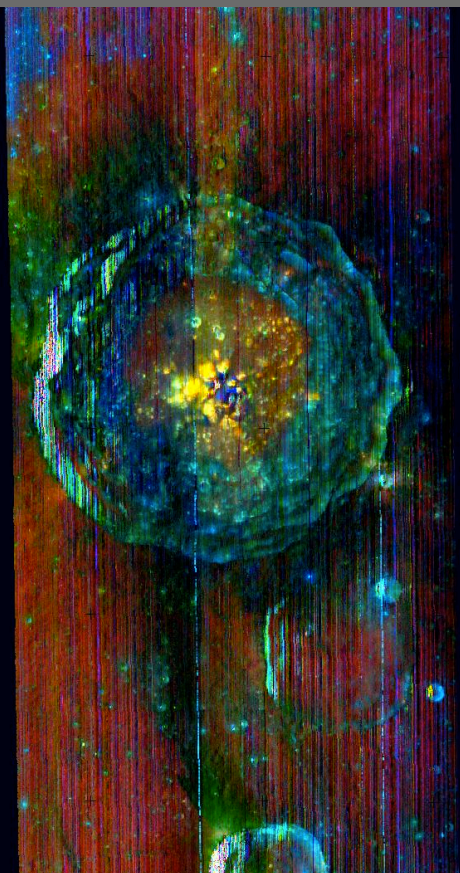
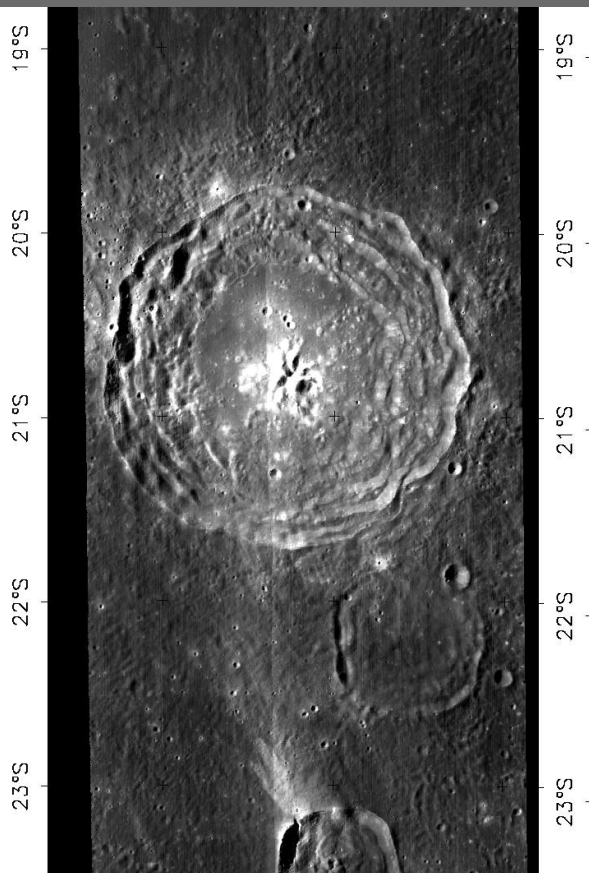
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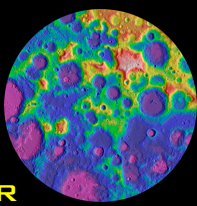
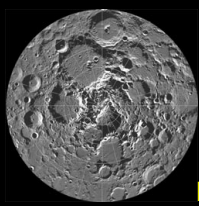


Crater Mineralogy

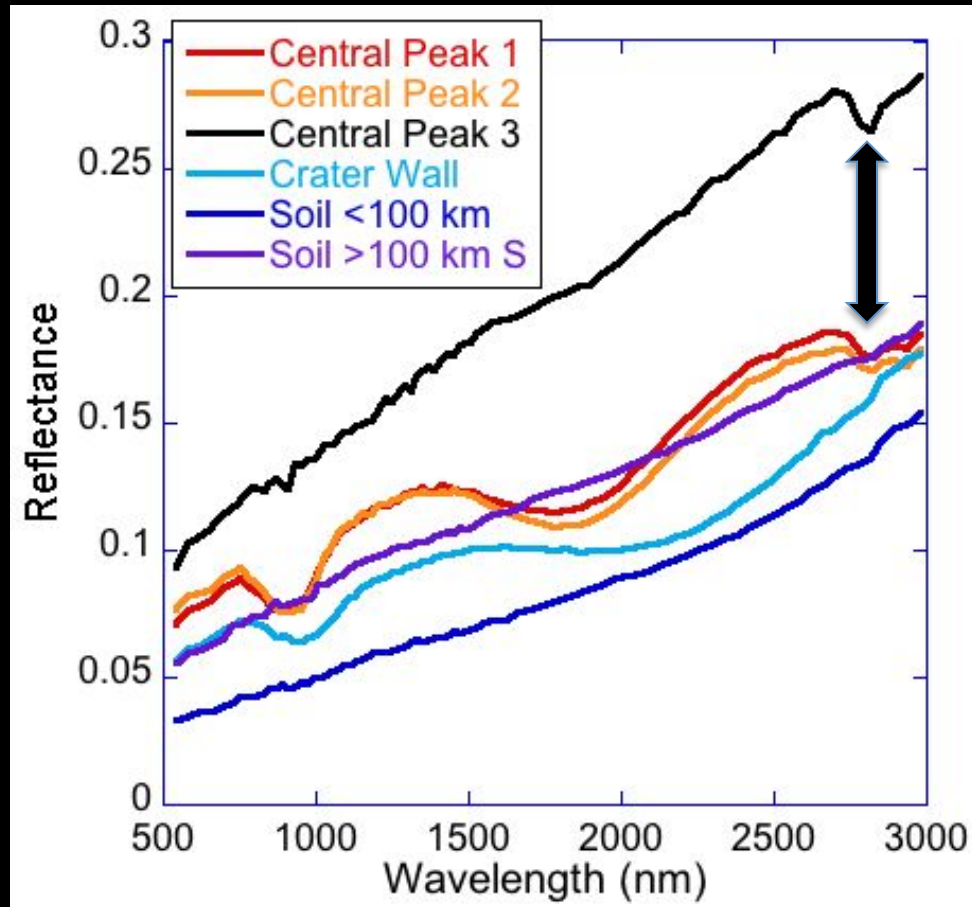
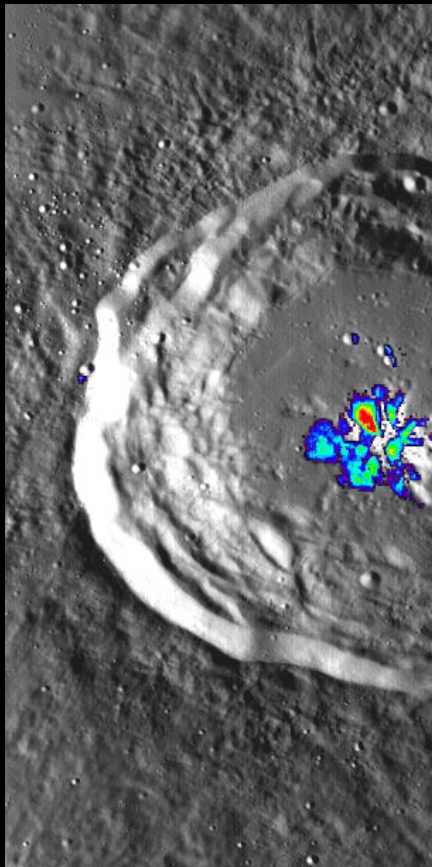


Yellow: Low-calcium pyroxene Black: Anorthosite
Blue/Cyan/Purple: Mare basalt or enriched in high-calcium pyroxene

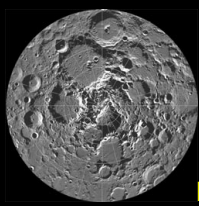




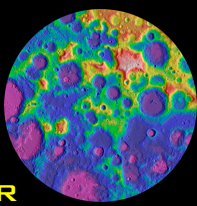
Hydroxyl at Bullialdus crater



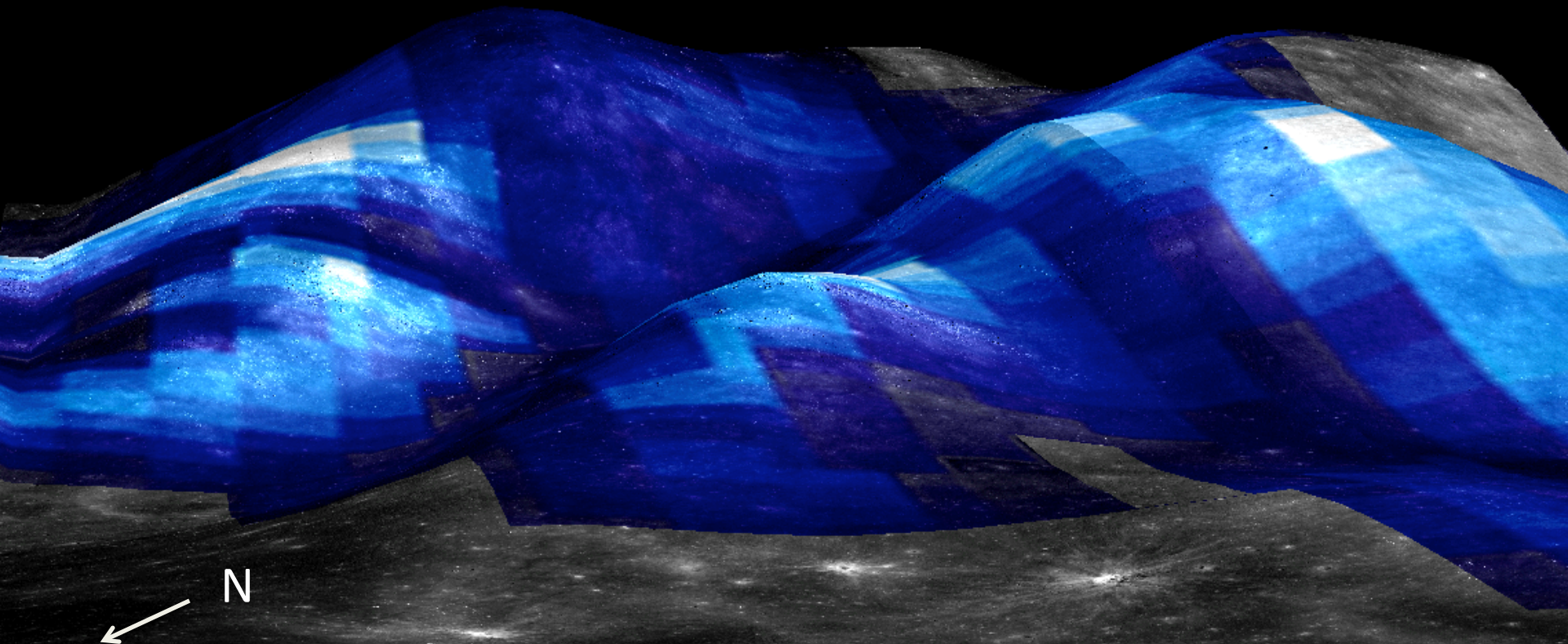
Absorption at 2.8 μm : OH^- strongly enhanced *only in central peak*



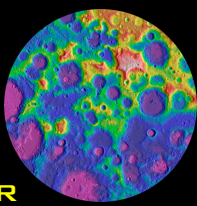
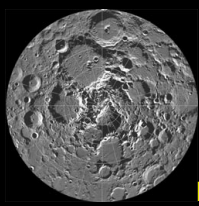
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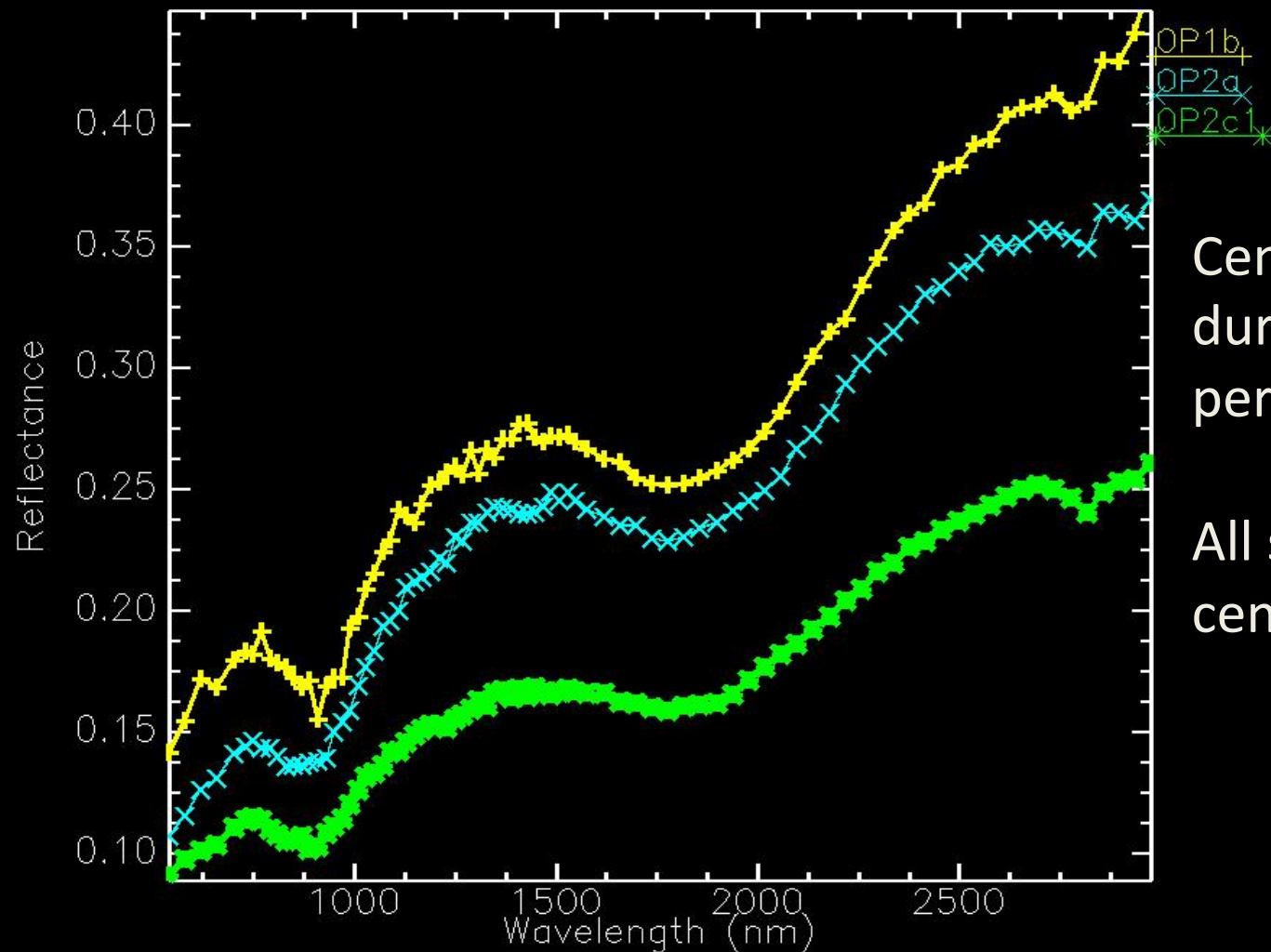
Hydroxyl overlain on NAC + LROC DEM



View is from North. Color ramp grades from dark blue (low) through white (high) OH^- abundance. High OH^- appears to correlate with higher albedo, more boulder-rich regions of the central peak.

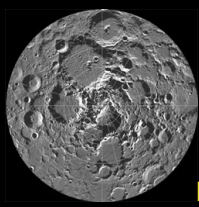


Hydroxyl at Bullialdus crater



Central peak imaged
during three optical
periods.

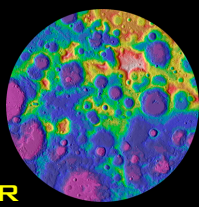
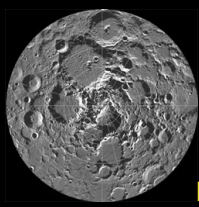
All show OH⁻ band in
central peak



Hydroxyl at Bullialdus crater



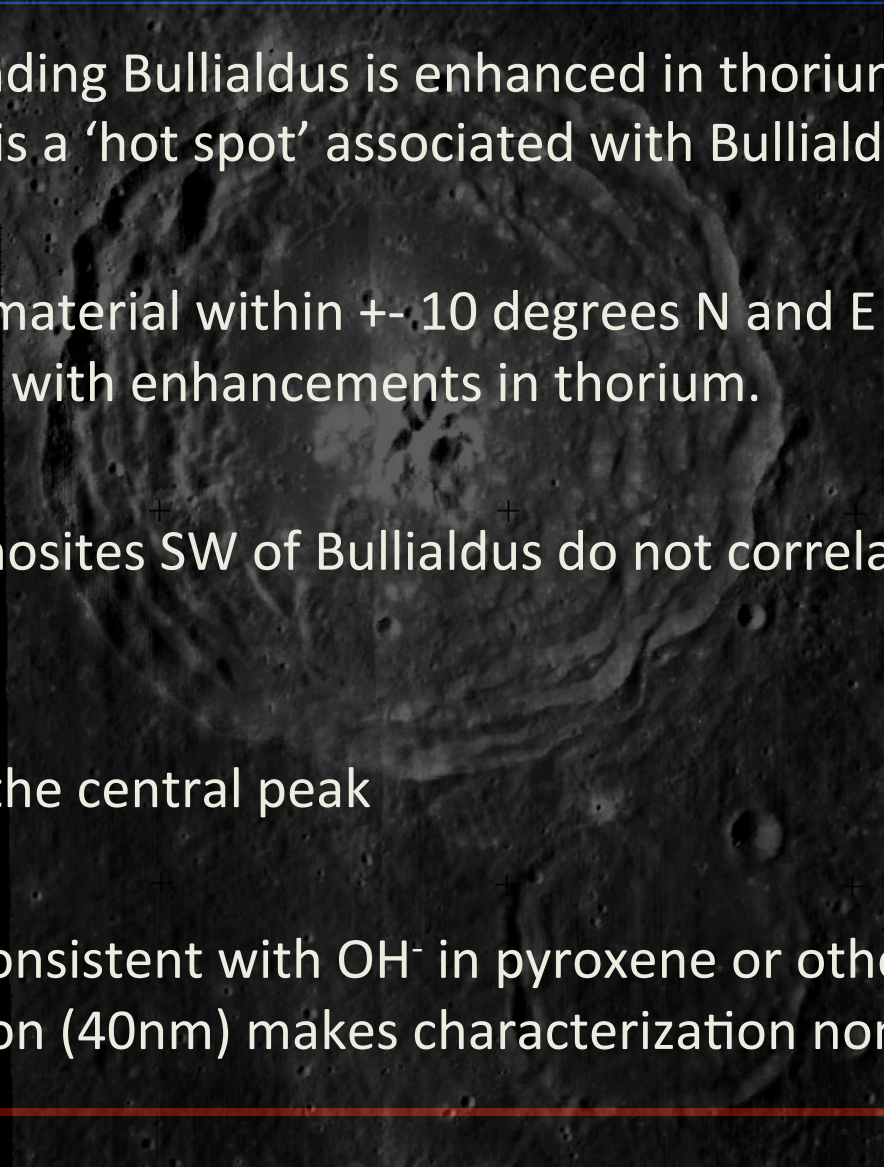
- Typically, non-polar OH⁻ is likely to be present as a thin surficial layer, likely produced by interactions of the solar wind with the lunar regolith (e.g., Pieters et al., 2009; Sunshine et al., 2009).
- Bullialdus was fully imaged by M³ at three different times in the lunar day. The 2.8 um absorption is observed and is of roughly equivalent strength during all of them.
- Highlands soils typically exhibit stronger OH⁻ bands, but those surrounding Bullialdus show little to no absorption.
- Fresh craters are observed to exhibit stronger OH⁻ bands, potentially due to an abundance of fractured bonds facilitating in-situ OH⁻ production. Bullialdus central peak is relatively immature due to mass wasting, but not freshly impacted.

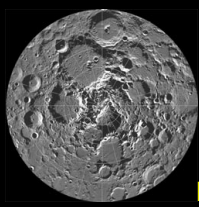


Bullialdus Crater: KREEP, Norite and OH⁻



- While the region surrounding Bullialdus is enhanced in thorium relative to the bulk Moon, there is a 'hot spot' associated with Bullialdus.
- Noritic and anorthositic material within ± 10 degrees N and E of Bullialdus also correlates with enhancements in thorium.
- 'Normal' highland anorthosites SW of Bullialdus do not correlate with enhanced thorium.
- OH⁻ is enhanced only in the central peak
- Position of OH⁻ band is consistent with OH⁻ in pyroxene or other silicates, but low spectral resolution (40nm) makes characterization non-unique

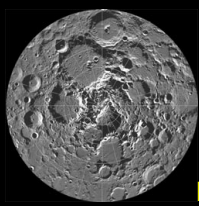




Bullialdus Crater: Implications



- The mineralogy of the central peak of Bullialdus is consistent with material from a KREEP-rich, Mg- or Alkali-suite pluton. Deconvolved thorium abundance is most consistent with the Alkali suite. If the OH^- is internal to the rocks, it supports OH^- enrichment in late stage urKREEP liquids.
- As a mafic pluton, less OH^- would have degassed than in KREEP-rich basalts.
- Further work is ongoing to characterize the nature of the OH^- band in more detail and to place bounds on the amount of OH^- detected.

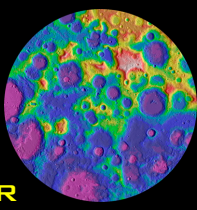
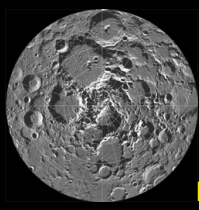


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Bullialdus Crater



- 20.7°S, 22.2°W on western edge of Mare Nubium
- 61 km diameter
- Eratosthenian-aged
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- Layered mafic pluton (Pieters, 1991)
- Other possibilities include impact through thin old basalt flows, or through a differentiated melt sheet from the Nubium impact (Tompkins et al., 1994).
- Radiative transfer modeling suggests immature regions of the peak range from anorthositic norite, anorthositic gabbro-norite to norite (Cahill and Lucey, 2007),
- Norites in central peak modeled to be $>Mg_{75}$ (Klima et al. 2011).